

Mon.	5.5 -.7 Curving Motion	RE 5.b
Tues.		EP 5, HW5: Ch5 Pr's 16, 19, 45, 48, 64(c&d)
Wed.	6.1-.4 (.21) Introducing Energy & Work Quiz 5	
Fri.	3pm – Visit from Columbia Rep	

Ch. 5 – Rate of change (if any) of momentum

$$\sum_{all} \vec{F}_{\rightarrow system} = \frac{d\vec{p}}{dt}$$

Special Cases

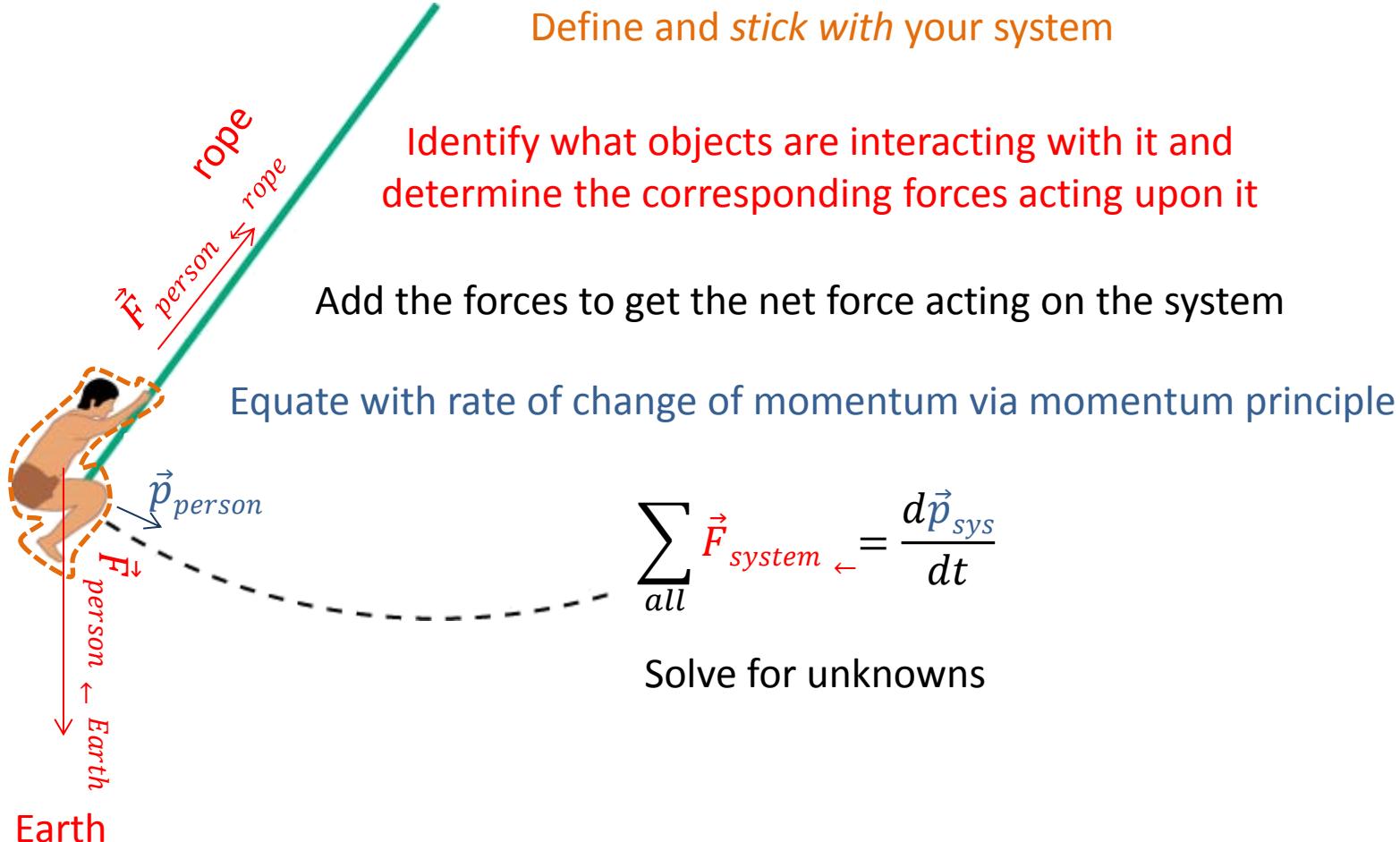
Equilibrium

$$\sum_{all} \vec{F}_{\rightarrow system} = \frac{d\vec{p}}{dt} = 0$$

Uniform Circular Motion

$$\frac{d|\vec{p}|}{dt} = 0$$

The process for force problems



Changing Momentum: Magnitude and Direction

$$\vec{F}_{net \rightarrow obj} = \frac{d\vec{p}}{dt} = \frac{d(\|\vec{p}\|\hat{p})}{dt} = \underbrace{\frac{d\|\vec{p}\|}{dt}\hat{p}}_{\text{Speeding/slowing}} + \underbrace{\|\vec{p}\|\frac{d\hat{p}}{dt}}_{\text{changing direction}}$$

Should point... Parallel to
momentum vector Parallel to
momentum vector

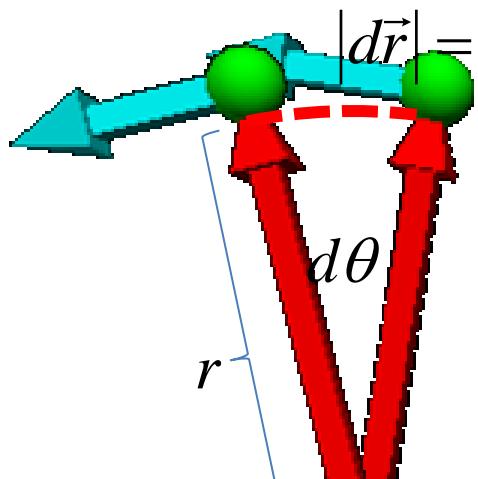
Special Case: Uniform Circular Motion (only direction changing)

$$\frac{d\vec{p}}{dt} = \frac{d(\|\vec{p}\|\hat{p})}{dt} = \frac{d\|\vec{p}\|}{dt}\hat{p} + \|\vec{p}\|\frac{d\hat{p}}{dt}$$

similarly

$$\vec{v} = \frac{d\vec{r}}{dt} = \frac{d\|\vec{r}\|}{dt}\hat{r} + \|\vec{r}\|\frac{d\hat{r}}{dt}$$

comparing



$$v = \left| \frac{d\vec{r}}{dt} \right| = \left| \frac{rd\theta}{dt} \right|$$

$$\left| \frac{d\vec{r}}{dt} \right| = \left| \vec{r} \right| \left| \frac{d\theta}{dt} \right|$$

$$\left| \frac{d\hat{r}}{dt} \right| = \left| \frac{d\theta}{dt} \right| = \frac{1}{\left| \vec{r} \right|} v$$

Rate of change of
position vector's
direction

Changing Momentum: Magnitude and Direction

$$\vec{F}_{net \rightarrow obj} = \frac{d\vec{p}}{dt} = \frac{d(\|\vec{p}\|\hat{p})}{dt} = \underbrace{\frac{d\|\vec{p}\|}{dt}}_{\text{Speeding/slowing}} \hat{p} + \underbrace{\|\vec{p}\| \frac{d\hat{p}}{dt}}_{\text{changing direction}}$$

Should point... Parallel to
momentum vector

Speeding/slowing

changing direction

Perpendicular to
momentum vector



Special Case: Uniform Circular Motion

$$\left| \frac{d\hat{r}}{dt} \right| = \left| \frac{d\theta}{dt} \right| = \frac{1}{|\vec{r}|} v$$

Rate of change of position
vector's direction

Equals

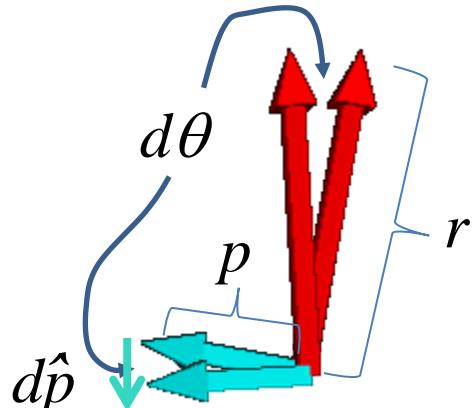
$$\frac{d\vec{p}}{dt} = \frac{d(\|\vec{p}\|\hat{p})}{dt} = \cancel{\frac{d\|\vec{p}\|}{dt}} \hat{p} + \|\vec{p}\| \frac{d\hat{p}}{dt}$$

Rate of change of momentum
vector's direction

$$\left| \frac{d\hat{r}}{dt} \right| = \left| \frac{d\hat{p}}{dt} \right| = \frac{1}{|\vec{r}|} v$$

direction?

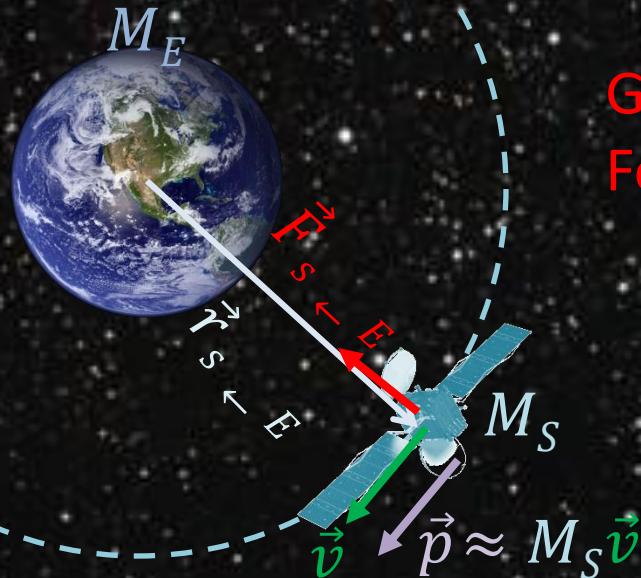
Opposite of r. $-\hat{r}$



See Vpython example

$$\frac{d\vec{p}}{dt} = -\|\vec{p}\| \frac{v}{|\vec{r}|} \hat{r}$$

Application: Circular Gravitational Orbits



System: Satellite

Gravitational Force

$$\vec{F}_{\text{net}} = \frac{d\vec{p}}{dt}$$

$$G \frac{M_E M_S}{|r_{s \leftarrow E}|} \hat{r}_{s \leftarrow E} = -|\vec{p}| \frac{|\vec{v}|}{|r_{s \leftarrow E}|} \hat{r}_{s \leftarrow E}$$

$$G \frac{M_E M_S}{|r_{s \leftarrow E}|} = |\vec{p}| |\vec{v}|$$

$$G \frac{M_E M_S}{|r_{s \leftarrow E}|} = M_S |\vec{v}| |\vec{v}|$$

$$G \frac{M_E}{|r_{s \leftarrow E}|} = \vec{v}^2$$

Circular Motion

Example: Geosynchronous Orbit

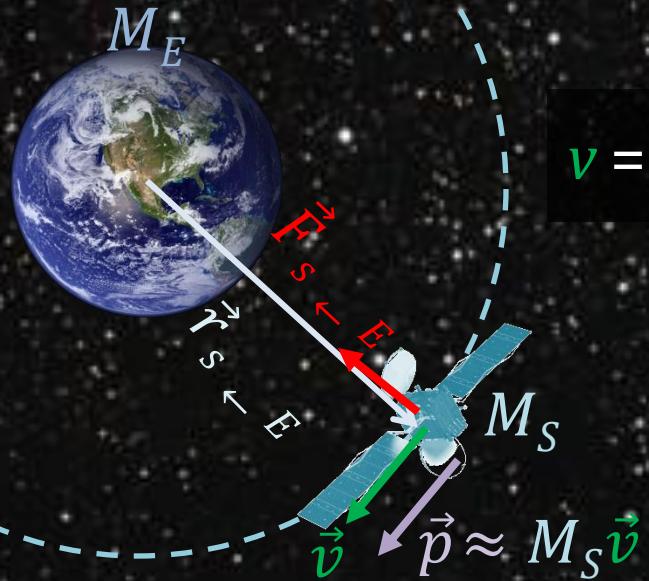
There's only one orbital radius for satellites that 'stay put' in the sky – orbit with the same period as the Earth spins: $T = 1$ day. What's the orbital radius?

$$v = \frac{\text{distance}}{\text{time}} = \frac{\text{Circumference}}{\text{Period}} = \frac{2\pi r_{s \leftarrow E}}{T}$$

$$G \frac{M_E}{|r_{s \leftarrow E}|} = \left(\frac{2\pi r_{s \leftarrow E}}{T} \right)^2$$

$$r_{s \leftarrow E} = \left(G M_E \left(\frac{T}{2\pi} \right)^2 \right)^{\frac{1}{3}} = \left(\left(6.7 \times 10^{-11} \frac{Nm^2}{kg^2} \right) (6 \times 10^{24} kg) \left(\frac{86,400 s}{2\pi} \right)^2 \right)^{\frac{1}{3}} = 4.2 \times 10^7 m$$

Kepler's 3rd Law of Planetary Motion



$$v = \frac{\text{distance}}{\text{time}} = \frac{\text{Circumference}}{\text{Period}} = \frac{2\pi r_{s \leftarrow E}}{T}$$

$$G \frac{M_E}{|r_{s \leftarrow E}|} = v^2$$

$$G \frac{M_E}{|r_{s \leftarrow E}|} = \left(\frac{2\pi r_{s \leftarrow E}}{T} \right)^2$$

$$r_{s \leftarrow E} = \left(GM_E \left(\frac{T}{2\pi} \right)^2 \right)^{\frac{1}{3}}$$

$$(r_{s \leftarrow E})^3 = GM_E \left(\frac{T}{2\pi} \right)^2$$

$$(r_{s \leftarrow E})^3 = \frac{GM_E}{(2\pi)^2} T^2$$

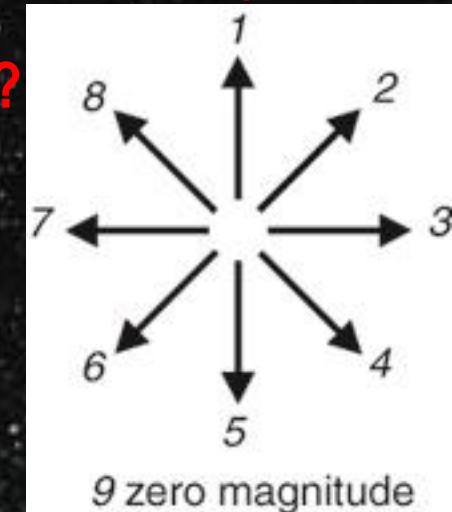
Orbital radius's relation to its period
Note: depends on mass of object orbited
not orbiting

Application: Circular Gravitational Orbits

The Moon travels in a nearly circular orbit around the Earth, at nearly constant speed.



what is the direction of $\frac{d\vec{p}}{dt}$?

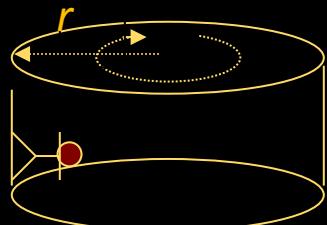


A geosynchronous satellite has an orbital radius of 4.2×10^7 m.

If the moon's period were 30 days (it's really about 27), what would be its orbital radius?

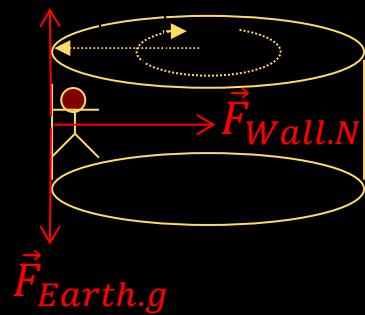
- a) 13×10^7 m
- b) 41×10^7 m
- c) 126×10^7 m
- d) 690×10^7 m

Circular Motion Example: Say you had a 1 km radius space-station, with what period must it spin to provide a normal force equal to that on Earth?

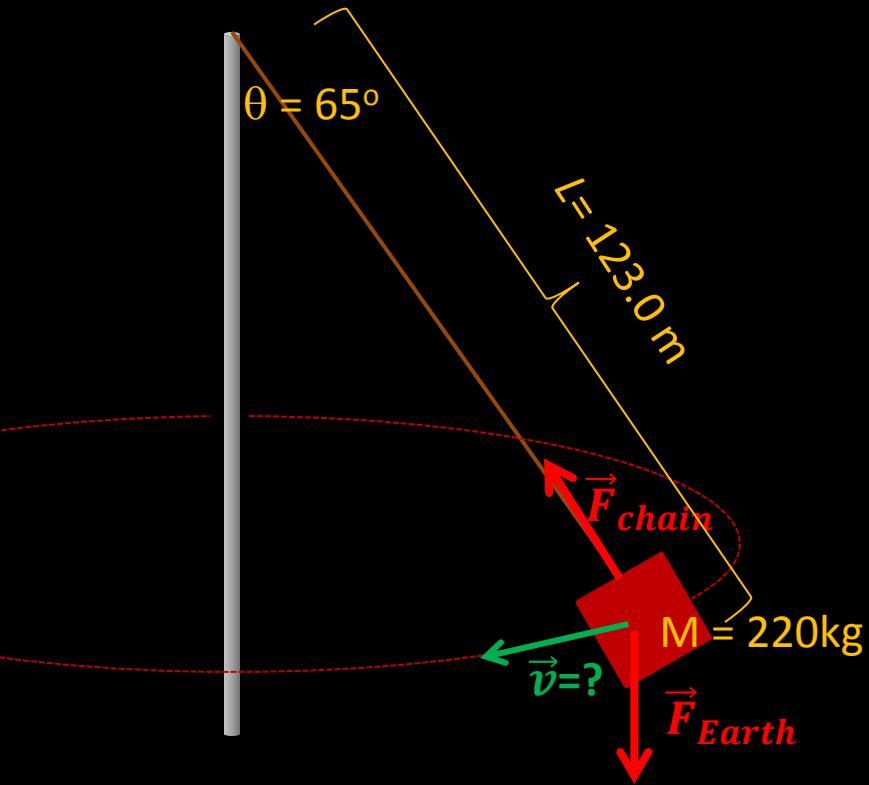


Homework: spinning carnival ride – friction holds you up

$$\vec{F}_{Wall.f} \leq \mu_s \vec{F}_{Wall.N}$$



(Know and care what interactions provide net force) Person rides on a “swing” ride at a carnival. The man rides in a swing at the end of a 12.0 m chain which hangs out at 65° up from vertical. Assuming uniform circular motion, and a combined, chair+rider, mass of 220 kg, (a) what is the tension in the chain? (b) what is tangential speed of the chair?



Circular Motion: Banked Curves

$$\vec{F}_{net} = \frac{d\vec{p}}{dt}$$

x-component of forces and
change in momentum

$$\frac{mg}{\cos(\theta)} \sin(\theta) = m \frac{v^2}{|r|}$$

$$\hat{x}: F_{c \leftarrow R.x} = |\vec{p}| \frac{|\vec{v}|}{|r|}$$

$$\vec{p} \approx m\vec{v}$$

$$F_{c \leftarrow R.x} = m \frac{v^2}{|r|}$$

$$F_{c \leftarrow R} \sin(\theta) = m \frac{v^2}{|r|}$$

y-component of forces and
change in momentum

$$\hat{y}: F_{c \leftarrow E} + F_{c \leftarrow R.y} = 0$$

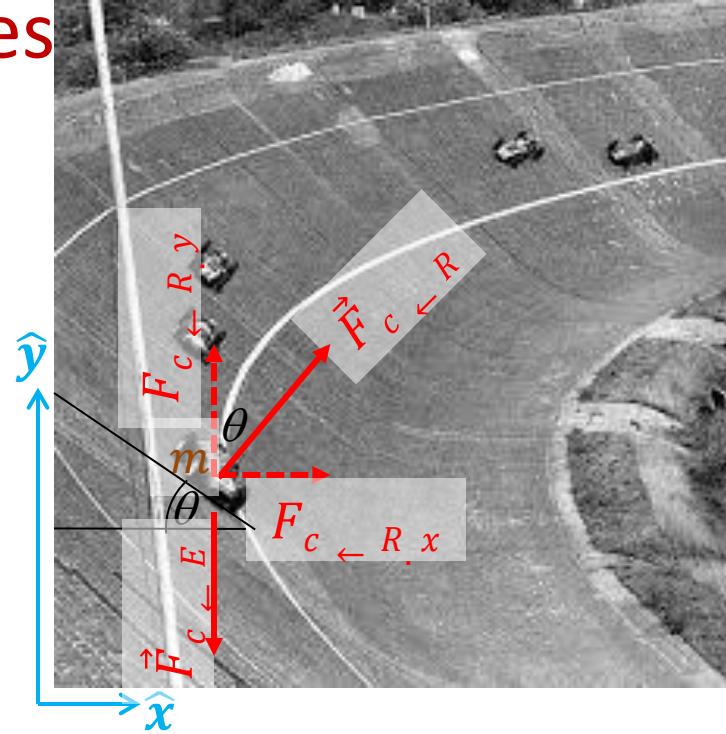
$$-mg + F_{c \leftarrow R.y} = 0$$

$$F_{c \leftarrow R} \cos(\theta) = mg$$

$$F_{c \leftarrow R} = \frac{mg}{\cos(\theta)}$$

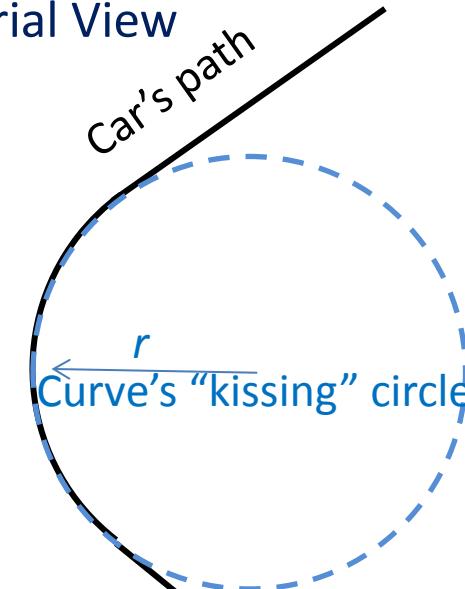
For a given radius of curvature and expected speed, there's a particular banking angle.

$$g \tan(\theta) = \frac{v^2}{|r|}$$



Why is the track banked?

Aerial View

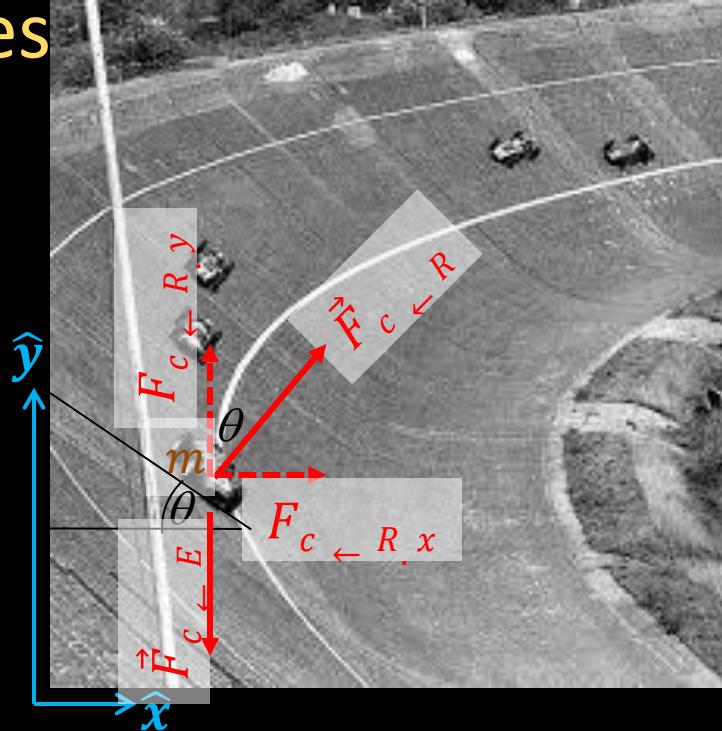


Circular Motion: Banked Curves

$$g \tan(\theta) = \frac{v^2}{|r|}$$

For a given radius of curvature and expected speed, there's a particular banking angle.

If you're taking the interchange between I-10 West and I-125 South, you may well be going 65 mph around a curve with a radius of curvature of 1/8 mi. If it was designed for this speed, at what angle should it be banked?



If your tires have a coefficient of static friction of 0.9, how fast could you take the curve without sliding sideways?

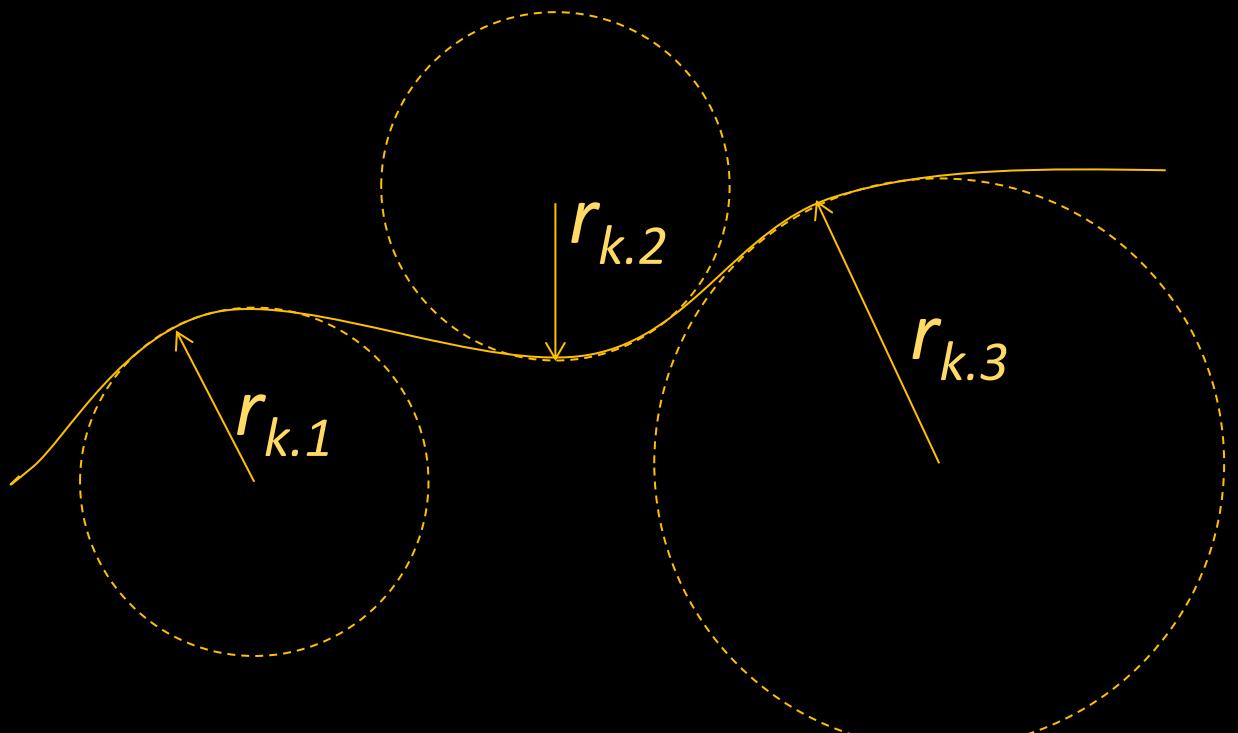
Circular Motion: Banked Curves

Why do planes bank when they turn?



Circular Motion: Generalized – “Kissing” Circles

$$\vec{F}_{net \rightarrow obj} = \frac{d(|\vec{p}| \hat{p})}{dt} = \frac{d(|\vec{p}|)}{dt} \hat{p} + |\vec{p}| \frac{d(\hat{p})}{dt}$$



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